

# Terrestrial Ground Penetrating Radar Studies as a Means of Constraining Lunar Near-Surface Properties

P. Russell<sup>1</sup>, J. Grant<sup>1</sup>,  
K. Williams<sup>2</sup>, B. Bussey<sup>3</sup>

1 Center for Earth and Planetary Studies, Smithsonian Institution

2 Dept. Earth Sciences and Earth Science Education, Buffalo State College (SUNY)

3 Planetary Exploration Group, JHU/APL

# Why use GPR?

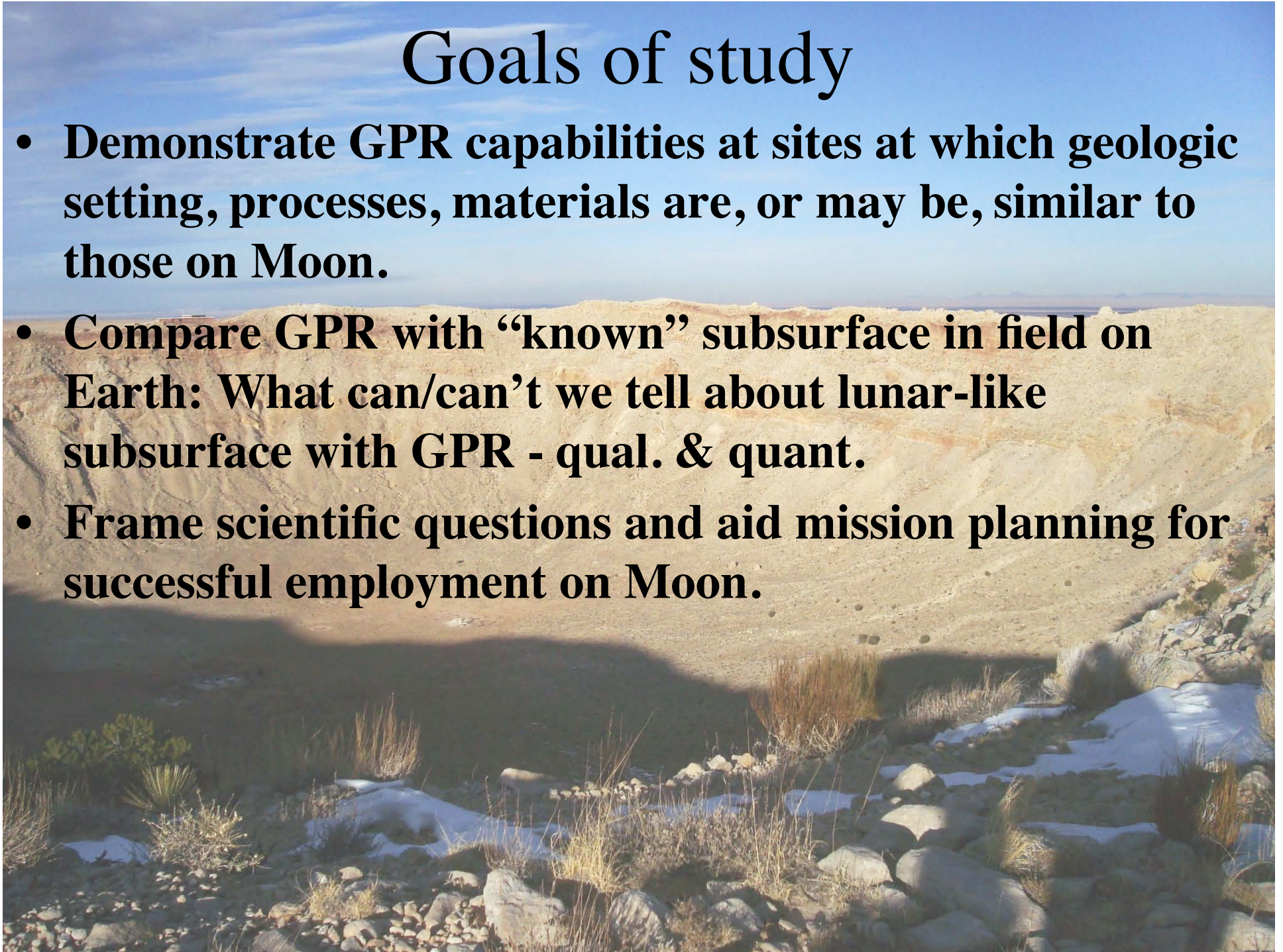
- Vertical, subsurface dimension is key to understanding local geologic processes over time.
- Location and physical properties of subsurface materials important for engineering.
- Complementary to drilling.
  - GPR: spatial coverage, deployment & use
  - Drill: materials analysis

# What can we see?

- Reflection off boundaries of differing dielectric properties (speed of waves in medium).
- Dielectric constant, density, degradation / weathering / fracturing, moisture, clay.
- 100 - 500 MHz to 0.5 m - ~8 m depth on Earth, same or more on Moon.
- Planar interface: Layering - eg lava flows, cinders, ejecta.
- Point reflector: Blocks / clasts - eg ejecta, fractured rock/basement.
- Voids - eg lava tubes

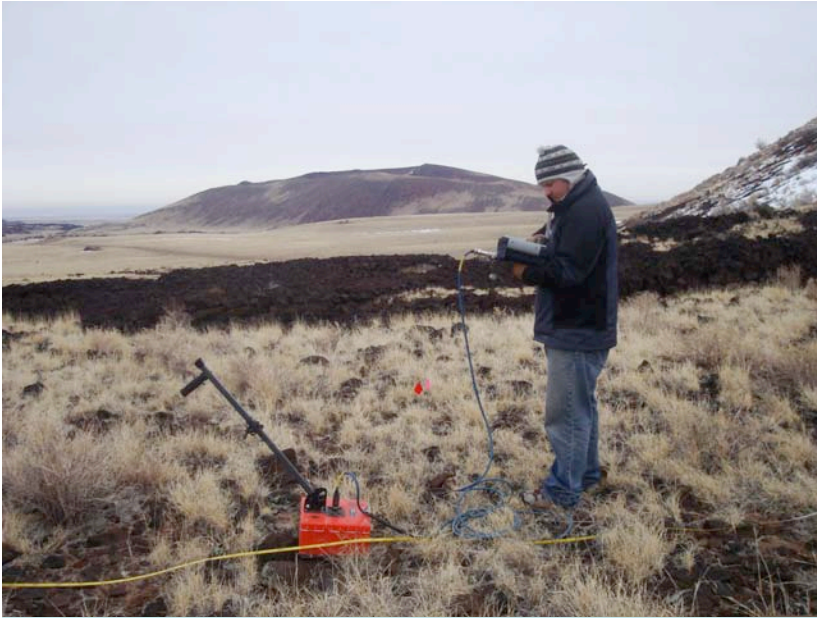
# Goals of study

- **Demonstrate GPR capabilities at sites at which geologic setting, processes, materials are, or may be, similar to those on Moon.**
- **Compare GPR with “known” subsurface in field on Earth: What can/can’t we tell about lunar-like subsurface with GPR - qual. & quant.**
- **Frame scientific questions and aid mission planning for successful employment on Moon.**



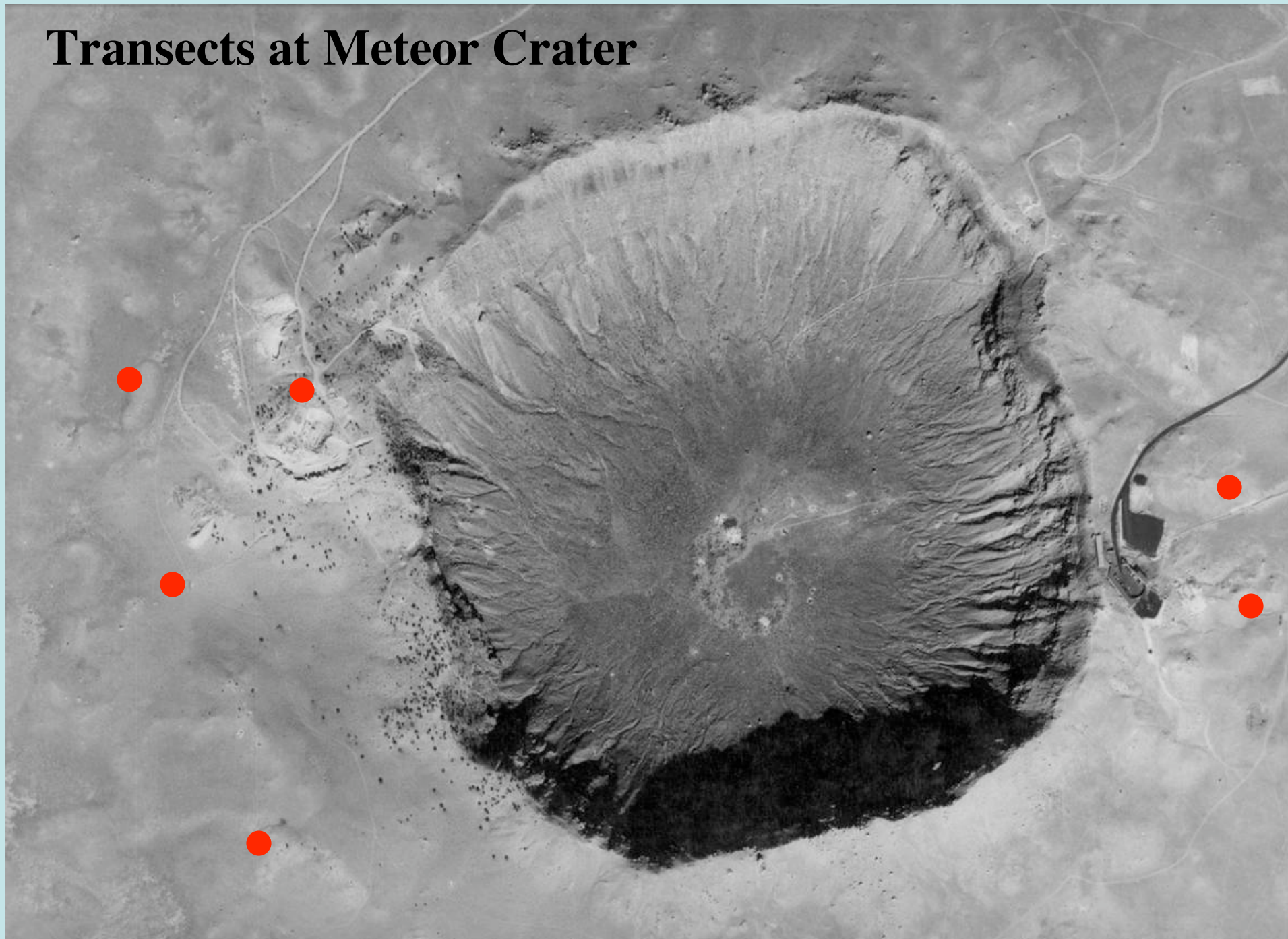


# Field work



- 200 MHz & 400 MHz transceivers.
  - GSSI, not prototype - focus on science
- Barringer Meteor Crater: Block-size distribution in ejecta.
- SP cone & lava flow: Layering in cinders and emergence of lava flow at base of cone.
- (Sunset Crater: Cinder-lava contact and layering.)

# Transects at Meteor Crater





# Meteor Crater Ejecta Lobes

<- to crater

ejecta

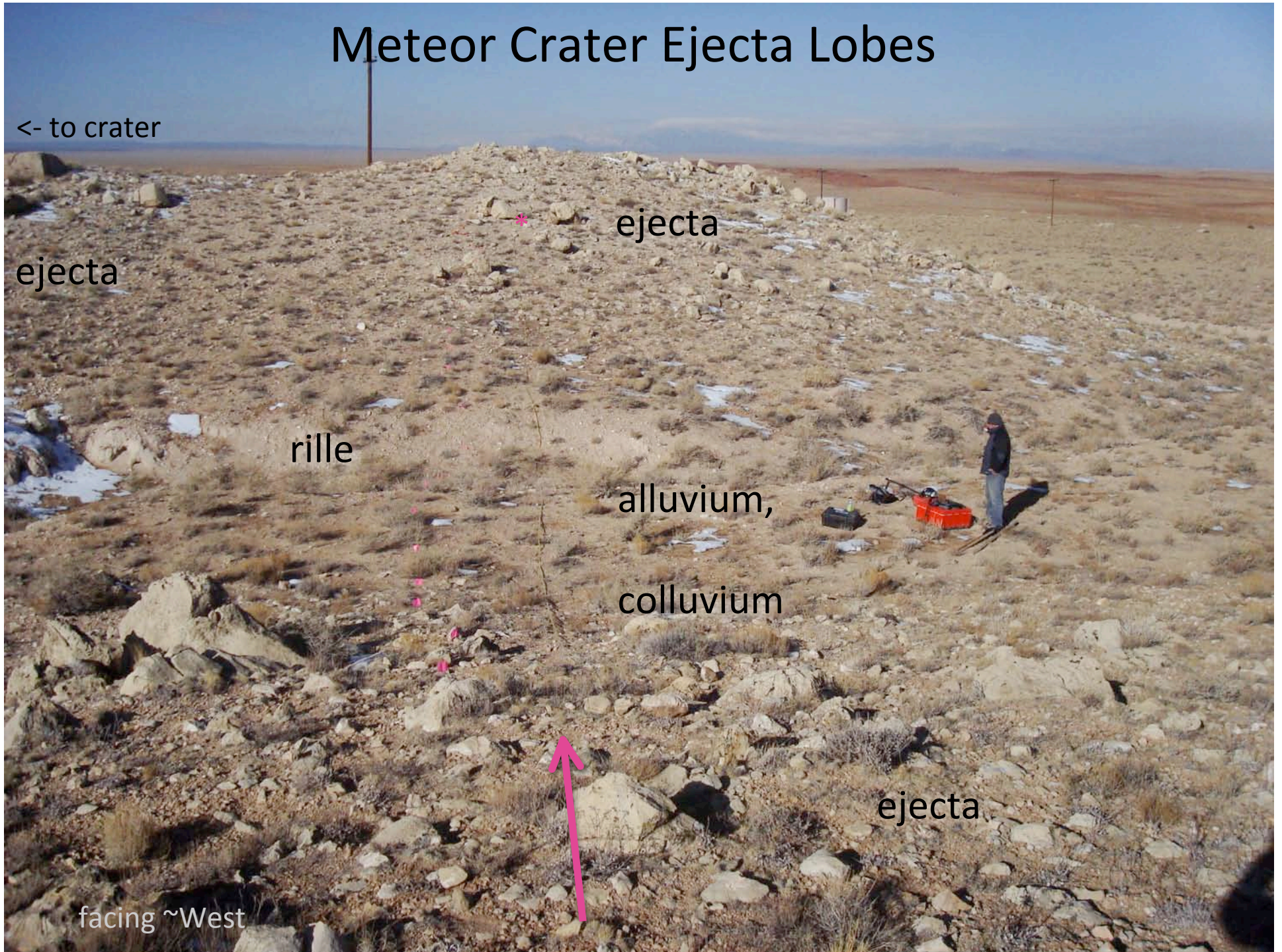
ejecta

rille

alluvium,  
colluvium

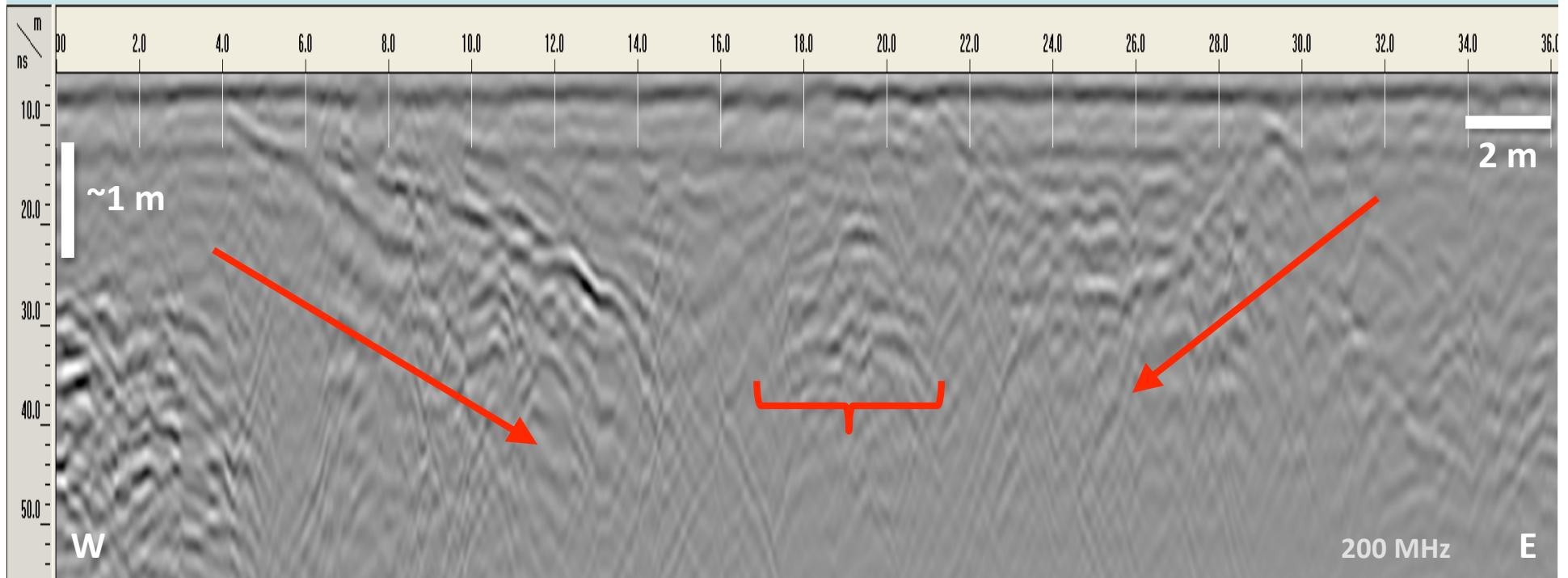
ejecta

facing ~West





# Meteor Crater Ejecta Lobes



ejecta surface dipping  
under co/alluvium

subsurface reflector rising  
relative to rille surface

ejecta surface dipping  
under co/alluvium



# Meteor Crater Ejecta Blocks

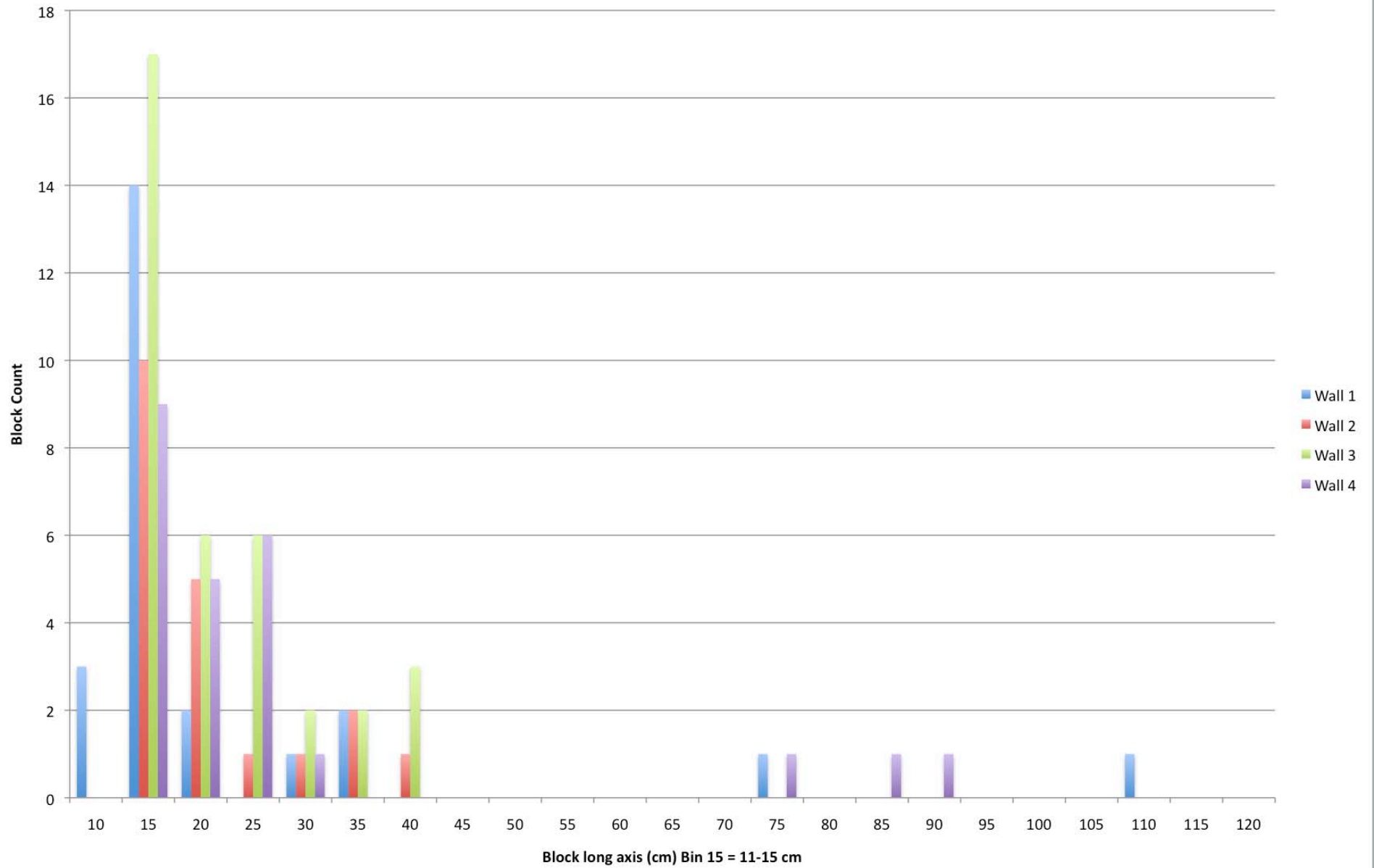




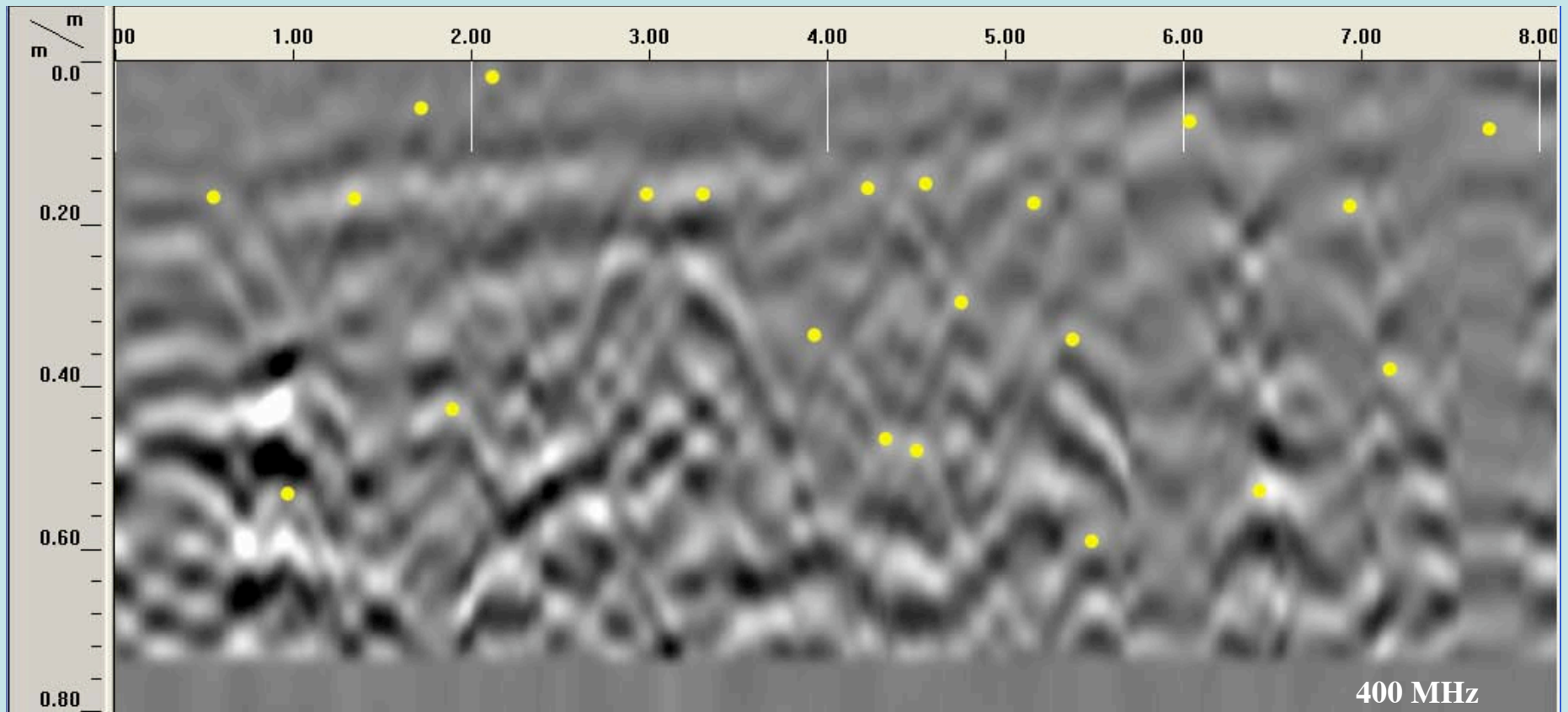




## Block Size-Frequency



# GPR



**1 m ~ 20 ns from metal plate calibration.**

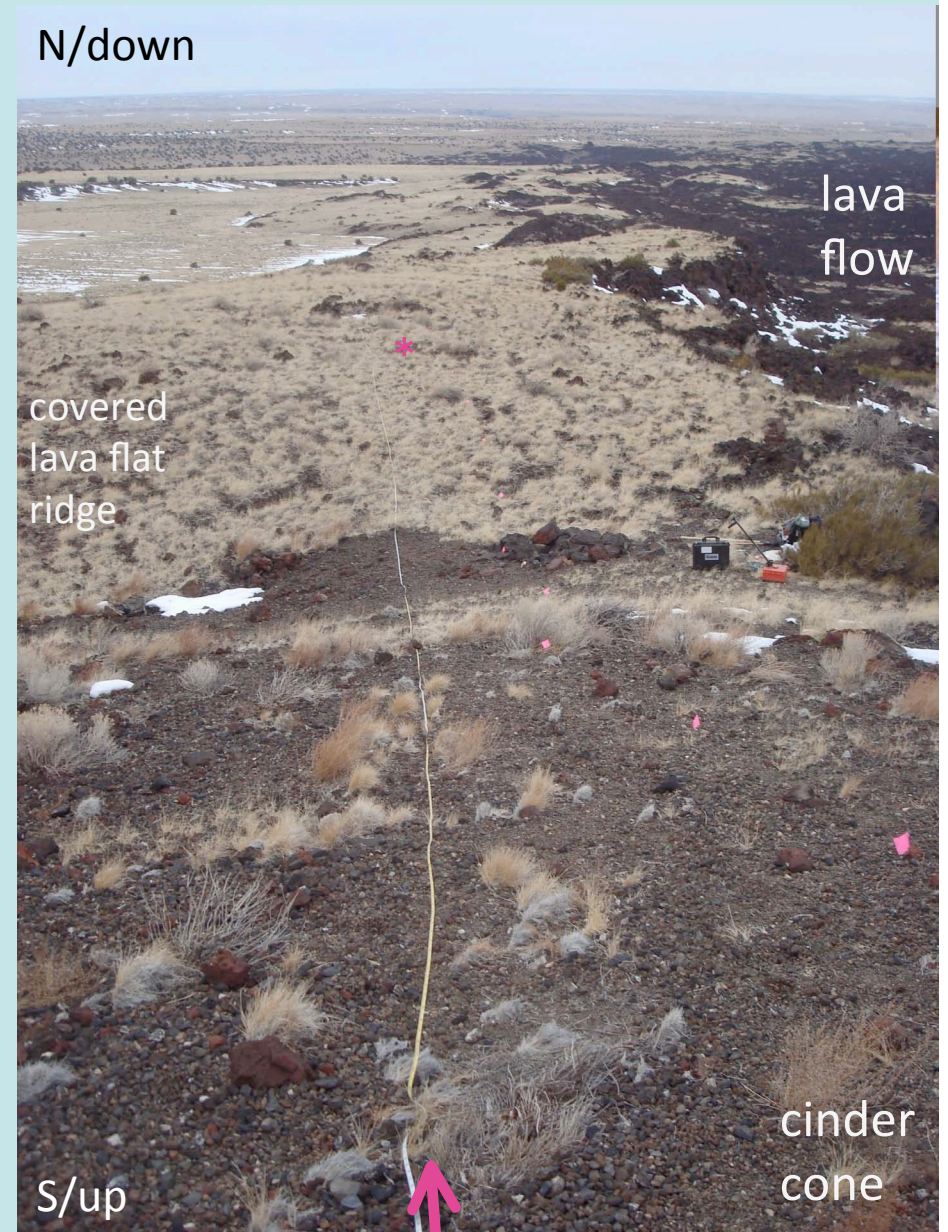
**Both GPR and wall photos reveal ~11 blocks per m<sup>3</sup>.**

**Decrease in number of blocks w depth.**

**200 MHz sees deeper, but fewer/larger blocks.**

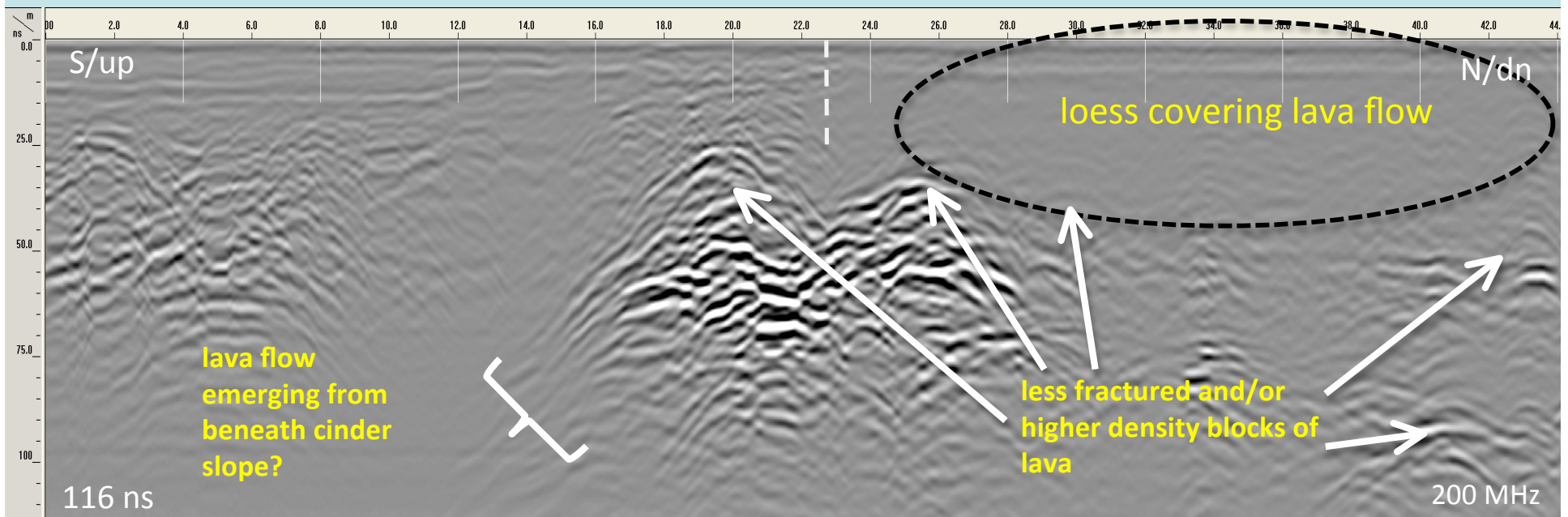
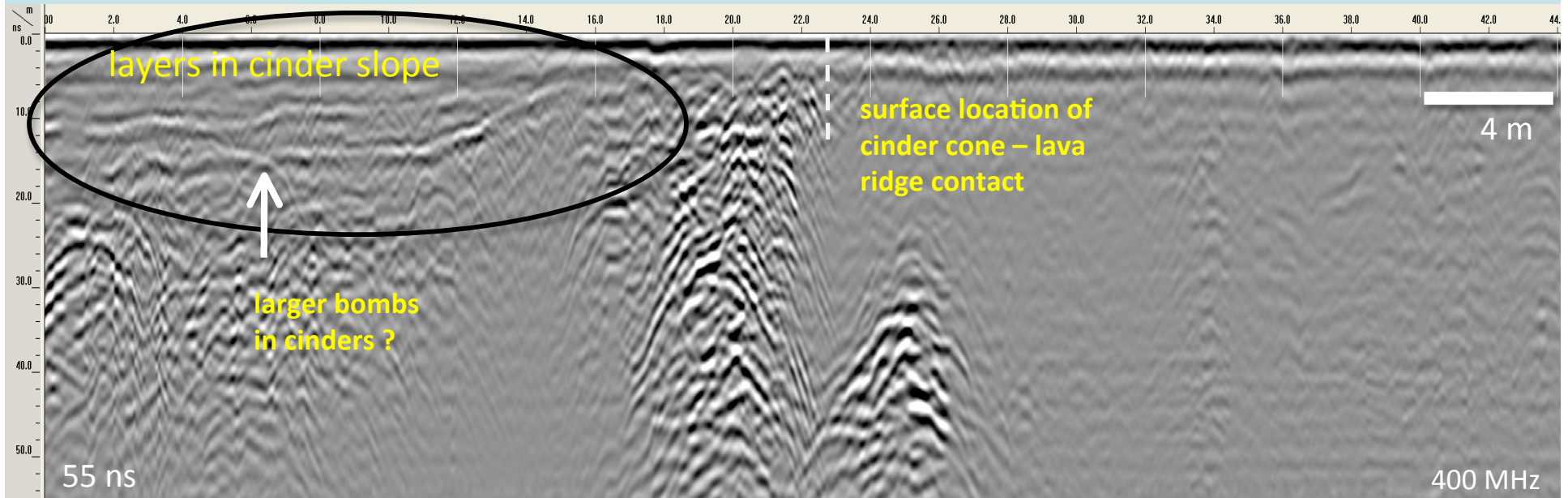


# SP Cinder Cone & Lava Flow





# SP Cinder Cone & Lava Flow





# Implications

- GPR can identify useful geologic information in the subsurface.
- Penetration likely greater on Moon.
- Calibration needed to convert to depth.
- Rover deployment: in contact with ground or elevated above.
- Complement orbiting radar.